5. THERMAL CONVERSION OF SOLAR RADIATION

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5.1 Introduction

The conversion of solar radiation into thermal energy happens in nature by absorption in the earth surface, planetary ocean and vegetation. Solar collectors are utilized to control this absorption and direct the thermal energy to consumers. This energy is conveyed from collectors to consumer by a liquid or gaseous heat carrier. The main parameter of this heat carrier is the temperature when leaving the collector, which depends on:
- power density of the incident radiation,
- concentration coefficient of the radiation
- sun position tracking
- optical efficiency of the collector,
- thermal efficiency of the collector.

Photovoltaic panels may be connected in series to obtain a higher voltage or in parallel to obtain a higher current intensity. On the contrary, a higher temperature of the heat carrier cannot be obtained by series connection of some collectors, because the greater is the number of connected collectors, higher are the thermal losses. For this reason many different types of collectors were built, each of them having a specific field of interest.

A general classification of solar collectors can be made by the concentration factor point of view:

$$c = \frac{S_C}{S_R}$$

where $S_C$ represents the entire area of the collector exposed to the sun, and $S_R$ is the heated area of the collector.

In this respect, there are two categories:
- collectors without concentration of the solar radiation, $c = 1$
- collectors with concentration, $c > 1$.

The first type of collector is intended for heat carrier temperature up to 100°С, while the second type collectors can reach a temperature level up to 3000-5000°С.

5.2 Collectors without concentration

Because $c = 1$, $S_C = S_R$ or, with other words, the exposed surface and the heated surface are equal. This type of collector presents some advantages:
- both components of solar radiation, the direct and the diffuse ones are utilized;
- an accurate tracking of the sun path is not compulsory;
- simple structure, easy available materials;
- small investment and maintenance expenses;
- great number of consumers for heat at low temperatures;
- good efficiency by comparison with the photovoltaic conversion.

The simplest type of collector without concentration is the flat collector (fig.1). Its operation is based on the so-called “greenhouse effect” which consists in selective behavior of transparent bodies crossed by radiations with different wave lengths. Thus, the solar radiation, having short wave length, crosses glass and air without important attenuation while the infrared radiation emitted by the receiving plate, having much greater wave length, is strongly attenuated.

After crossing the colorless glass, the solar radiation is absorbed in the receiving plate, black painted. The plate warms up until an equilibrium is established between the entering solar energy and the heat losses.

**5.2.1 Optical efficiency of the flat collector**

The optical efficiency is defined as the ratio between the absorbed energy in the collector plate and the solar radiating power applied on the upper surface of the collector. The difference between these quantities represents the losses owed to the reflexion, refraction and absorption in the transparent layers and as well as to the partial absorption of the light in the receiving plate.

Taking into account only one transparent layer and neglecting, for the beginning, the absorption of the light in this layer, the transmission coefficient, $D_R$ [p.u.] owed to the reflexion and refraction processes can be determined (fig. 2).
The transmitted energy through the transparent layer can be evaluated by adding the successive components in fig.2

\[ D_R = (1 - R)^2 + R^2 (1 - R)^2 + \ldots + R^{2n} (1 - R)^2 = (1 - R)^2 \left( 1 + R^2 + \ldots + R^{2n} \right) \]

By adding the terms between brackets, results

\[ D_R = (1 - R)^2 \frac{1 - R^{2(n+1)}}{1 - R^2} \quad (2) \]

Because \( R < 1 \), \( R^{2(n+1)} \to 0 \), so that

\[ D_R = \frac{1 - R}{1 + R} \quad (3) \]

If there are \( s \) transparent layers, after a similar reasoning, the transmission coefficient becomes

\[ D_R = \frac{1 - R}{1 + (2s - 1)R} \quad (4) \]

The reflexion coefficient \( R \) can be evaluated by the Fresnel equation for the reflexion process of an un-polarized radiation which propagates from a medium characterized by a refraction index \( n_1 \) into another medium with refraction index \( n_2 \):

\[ R = \frac{1}{2} \left[ \frac{\sin^2(\theta_2 - \theta_1)}{\sin^2(\theta_2 + \theta_1)} + \frac{\tan^2(\theta_2 - \theta_1)}{\tan^2(\theta_2 + \theta_1)} \right]. \quad (5) \]

\( \theta_1 \) and \( \theta_2 \) are the angle of incidence and respectively the angle of refraction, governed by the refraction law:

\[ n_1 \sin \theta_1 = n_2 \sin \theta_2. \]

If the incident radiation is perpendicular to the surface of the transparent layer,

\[ \theta_1 = \theta_2 = 0, \]

thus

\[ R = \left( \frac{n_1 - n_2}{n_1 + n_2} \right)^2. \quad (6) \]

The transmission coefficient \( D_R \) depends on the angle of incidence \( \theta_1 \) according the graph in fig. 3. An important conclusion resulting from this figure is that the optical efficiency of the flat collector maintains unaffected when the angle of incidence changes with \( \pm 40-45 \) degrees from the direct orientation to the sun.

The incident radiation is partially absorbed by the glass, so that the transmitted power, \( D \), depends on the path length through glass, \( d \), upon the exponential law:

\[ D = D_R e^{-kd}. \quad (7) \]

\( k \) is the attenuation coefficient, which depends on the chemical composition of the glass. For example, a glass with very low content of Fe has \( k = 0.04 \) cm\(^{-1}\), while a reprocessed glass can reach a value \( k = 0.32 \) cm\(^{-1}\).
Subtracting the transmitted power $D$ from $D_R$, the result is $A$ which represents the absorbed power having as consequence a warming of the glass:

$$A = 1 - e^{-kd} \quad (8)$$

Considering a glass of 0.5 cm thick, the light attenuation represents 7.5% for the best quality glass and 52% for the poorest quality one.

Further, the transmitted radiation is partially absorbed by the receiving plate, depending on the absorption coefficient $A_p$. A new repeated process of absorptions and emissions takes place, like that presented in fig. 5.

The incident radiation on the receiving plate will be considered as directional one, like it comes from the sun. The emitted radiation from the plate is diffuse, because the plate which must absorb the radiation as much as possible has a lusterless surface.

A new reflexion coefficient $R_d$ is introduced to evaluate the diffuse radiation reflected from the glass, whose value depends on the number of glass layers. For example, $R_d = 0.16$, 0.24, 0.29 if the number of glass layers, $s$, is 1, 2 or 3 respectively.

By adding the power components entering the receiving plate [u.r.], the result represents the optical efficiency of the collector.

$$\eta_{opt} = DA_p + DA_p (1 - A_p) + DA_p (1 - A_p)^2 R_d^2 + ... + DA_p (1 - A_p)^s R_d^n + ...$$

$$\eta_{opt} = DA_p [(1 - A_p) R_d + (1 - A_p)^2 R_d^2 + ... + (1 - A_p)^n R_d^n + ...]$$
Because $1 - A_p < 1$ and $R_d < 1$, the second term of the nominator can be neglected, so that the optical efficiency becomes

$$
\eta_{opt} = \frac{DA_p}{1 - (1 - A_p)R_d}.
$$

\[(9)\]

5.2.2 Thermal efficiency of the flat collector

The heat developed in the absorbing plate can be divided in three main components:
- the heat transported by the fluid to the consumer,
- the heat accumulated in the collector materials, which are warming up,
- the heat losses from collector to ambient space.

If the density of the solar radiation and the heat carrier mass flow are constant, the absorbing plate temperature is constant, too. As consequence, for this stationary operating mode, the accumulated heat in the collector materials can be neglected in the equation of energy balance.

The heat losses take place through all the outer surfaces of the collector. Because the collector structure differs on the upper direction by comparison with the others sides, the basic heat transfer mechanisms have different weight on these directions. On the upper side prevails the heat losses by radiation and convection. On the other sides of the collector, the conduction mechanism is dominant, because the temperature of the external surfaces is low enough to can neglect the radiation and convection losses.

To calculate the whole heat losses, the next equation may be considered:

$$
Q_p = K_t S(t_p - t_a),
$$

\[(11)\]

where
- $K_t$ is the whole heat transfer coefficient,
- $S$ is the area of the upper surface of the collector, equal with the receiving plate area
- $t_p$ is the average temperature of the receiving plate
- $t_a$ is the temperature of the surrounding air.

Because the global heat transfer coefficient is difficult to evaluate, an electrical similarity of the heat transfer can be suggested. Considering a potential difference $U_1 - U_2$ applied to a resistor $R$, the current intensity, may be written

$$
I = \frac{U_1 - U_2}{R}.
$$

Similarly, the heat transmitted, per unit of surface, between two parallel surfaces, having different temperatures $t_1 > t_2$ and the transfer coefficient $k$, is

$$
q = k(t_1 - t_2) = \frac{t_1 - t_2}{R_i}.
$$

Fig.5 The analogy between the electric current and the heat transfer
Thus, the thermal resistance is similar to the electric resistance, the temperature difference is similar to the electric voltage and the heat flow is similar to the electric current.

For example, the diagram in fig. 6 represents the equivalent electric circuit for the flat collector in fig. 1, having two transparent layers of glass. \( t_p \) is the temperature of the receiving plate, \( t_a \) is the air temperature, \( t_{g1}, t_{g2} \) are the temperatures of the two glass layers and \( t_c \) is the temperature on the side and back surfaces of the collector.

![Fig.6 The equivalent of the heat losses in the flat collector](image)

The electric analogy allows expressing the global thermal resistance, \( R_t \). On the upper side of the collector, the three thermal resistances are connected in series so that:

\[
R_{upp} = R_{t1} + R_{t2} + R_{t3}
\]

On the others sides of the collector, the thermal resistances are connected in series, too:

\[
R_b = R_{t4} + R_{t5}
\]

The global thermal resistance \( R_t \) will be:

\[
R_t = \frac{R_{upp} R_b}{R_{upp} + R_b},
\]

thus the global heat transfer coefficient, \( K_t \) is

\[
K_t = \frac{1}{R_t}.
\]

To calculate the heat resistances, the heat transfer processes must be analyzed. From theoretical point of view, the upper structure of the collector can be modeled by parallel surfaces having infinite sizes and different temperatures. Considering a pair of surfaces with the temperatures \( t_1 \) and \( t_2 \), the heat transfer between them, by radiation and convection, obeys the next equation:

\[
q_{12} = \alpha_C (t_1 - t_2) + \frac{\sigma (t_1^4 - t_2^4)}{\frac{1}{\varepsilon_1} + \frac{1}{\varepsilon_2} - 1}.
\] (12)

The first term in the right part of this equation represents the transfer by convection, while second term, by the radiation process. The significance of the symbols in the previous equation are:

- \( \alpha_C \) is the heat transfer coefficient by convection,
- \( \sigma \) is the Stefan-Boltzmann constant \((5.67 \times 10^{-8} \text{ W/m}^2 \text{K}^4)\),
- \( \varepsilon_1 \) and \( \varepsilon_2 \) are the emission coefficients by radiation from the two surfaces.

By extracting the term \((T_1 - T_2)\) as a common factor in this equation, results:
\[
q_{12} = \left[ \alpha_C + \frac{\sigma(T_1+T_2)(T_1^2+T_2^2)}{1+\frac{1}{\varepsilon_1}+\frac{1}{\varepsilon_2}+1} \right] (T_1-T_2) = (\alpha_C + \alpha_R)(T_1-T_2).
\]

This way, two heat resistances in fig. 1 can be written as:

\[
R_{t1} = \frac{1}{\alpha_{C1}+\alpha_{R1}}, R_{t2} = \frac{1}{\alpha_{C2}+\alpha_{R2}}.
\]

For the heat resistance \( R_{t3} \), the model structure is different, the atmosphere being one of the two surfaces. Thus, for the heat transfer by radiation to the atmosphere, the heat transfer coefficient is:

\[
\alpha_{R3} = \varepsilon_g \sigma \left( T_{g2} + T_o \right) \left( T_{g2}^2 + T_o^2 \right).
\]

For the heat transfer by convection to the atmosphere, an empirical equation is recommended:

\[
\alpha_c = 5.7 + 3.8v_t
\]

where \( v_t \) is the air speed parallels with the upper surface of the collector.

To apply this algorithm is necessary to know the temperatures of the parallel surfaces, that is difficult to obtain. As consequence, simpler methods are utilized for the collector designing:

- nomographs based on the analytic equations, valid for specific collector structure or materials;

- empirical equations, experimentally proved as valid for specific structure and materials.

A heat transfer by conduction produces through the thermal insulation on the back and side walls of the collector, according to the next equation:

\[
q_{is} = \frac{\lambda_{is}}{\delta_{is}} (t_p - t_a),
\]

where \( \delta_{is} \) represents the thermal insulation thickness, \( \lambda_{is} \) is the coefficient of thermal conduction of the insulating material.

Thus,

\[
R_{t4} = \frac{\delta_{is}}{\lambda_{is}}.
\]

The heat losses, by radiation and convection, from the side and back surfaces of the collector, can be neglected because the both temperatures are very close. As consequence, \( R_{t5} \to \infty \), respectively, \( k_s = 0 \).

### 5.2.3 The whole efficiency of the flat collector

In order to calculate the collector efficiency, some energy quantities must be defined:

- the incident energy on the upper surface of the collector

\[
W_1 = ESf,
\]

where \( E \) is the power density of the sun radiation, \( S \) is the area of upper surface of the collector and \( f \) is it’s orientation factor;

- the energy developed in the receiving plate
\[ W_2 = ESf\eta_{opt}; \]  
(18)

- the energy losses in the optical system of the collector

\[ W_3 = (1-\eta_{opt})ESf; \]  
(19)

- the whole heat losses

\[ W_4 = K_tS(t_p - t_a). \]  
(20)

Two possibilities to define the collector efficiency are relevant:

• the whole efficiency

\[ \eta_g = 1 - \frac{W_3 + W_4}{W_1}. \]

By replacing the above defined energy quantities, results

\[ \eta_g = \eta_{opt} - \frac{K_t(t_p - t_a)}{Ef}. \]  
(21)

• the inner (thermal) efficiency

\[ \eta_t = 1 - \frac{W_4}{W_2} = 1 - \frac{K_t(t_p - t_a)}{Ef\eta_{opt}}. \]  
(22)

These two efficiencies are linked by the equation

\[ \eta_g = \eta_{opt}\eta_t, \]  
(23)

so that \( \eta_g < \eta_t. \)

The collector efficiency depends on more factors like sun power density, collector orientation, the temperature of the receiving plate, the air temperature etc. For a good quality collector, in a normal regime of operation, the energy losses have the average values:

<table>
<thead>
<tr>
<th>Type of losses</th>
<th>Loss weight</th>
</tr>
</thead>
<tbody>
<tr>
<td>optical losses</td>
<td></td>
</tr>
<tr>
<td>by reflexion</td>
<td>2 x 4% = 8%</td>
</tr>
<tr>
<td>by refraction</td>
<td>2 x 3.5% = 7%</td>
</tr>
<tr>
<td>thermal losses</td>
<td></td>
</tr>
<tr>
<td>convection</td>
<td>17%</td>
</tr>
<tr>
<td>radiation</td>
<td>12%</td>
</tr>
<tr>
<td>conduction</td>
<td>3%</td>
</tr>
<tr>
<td>energy losses, total</td>
<td>47%</td>
</tr>
</tbody>
</table>

The global efficiency results to be 53%, much greater than that of the photoelectric cells. However, the temperature gap between the receiving plate and the air influences the whole efficiency, like fig. 7 shows. The nonlinear dependence is owed to the heat losses through radiation which depend on the 4th power of the temperature. The highest temperature of the plate is reached when the heat carrier does not circulate, while the collector is exposed to the sun. At this temperature all the energy losses from the collector equalize the received sun power. Is very important to know how great this temperature for an adequate designing of the collector is so that it could operate in every regime without damages.
An example of flat collector design is shown in fig. 8.

**5.2.4 The cylindrical collector**

A cylindrical collector consists in two tubes: the inner, metallic tube having a dark surface in order to absorb the solar radiation, surrounded by a transparent tube, fig. 9. Through the inner tube flows the heat carrier fluid.
This structure of collector presents some advantages by comparison with the flat collector:

- the illuminated surface area of the inner cylinder remains unaffected by the sun position on the sky;
- if the space between tubes is empty (advanced vacuum) the convection and conduction heat losses are avoided, so that the whole efficiency can be much improved.

![Different structures of cylindrical collectors](image)

Fig. 10 Different structures of cylindrical collectors

To increase the received sun power, the absorbing surface must increase too. However, this growth is limited because if the inner tube diameter grows, the volume of the fluid grows too, as well as the thermal inertia of the collector.

More cylindrical collectors can be connected in parallel and arranged on a plate surface like a flat collector. To utilize the solar radiation which passes between inner tubes, the surface of the outer tube can be covered with a mirror layer. Thus, the reflected radiation reaches the absorbing surface (fig. 10, left).

Another solution is to add an absorbing flat surface to the inner tube (fig. 10, right). In this manner, the entire solar radiation reaches the absorbing surface, but the independence of the sun position on the sky will be partially lost. However, remains unaffected the main advantage of heat losses diminution.

![Cylindrical collector design](image)

Fig. 11 Cylindrical collector design

Despite it is more expensive than the flat collector, the cylindrical one are nowadays much more asked because it’s better efficiency.

### 5.3 Solar radiation concentration

To obtain a concentration coefficient greater than 1, an optical system must be inserted between the sun and the absorbing surface (the receiver). This optical system is based either on the reflexion or refraction phenomena of the solar directional radiation. The diffuse radiation can not be concentrated.
Depending on the mirror geometry, the concentration coefficient may vary between 2 and $10^4$ that mean a similar growth of the power density on the absorbing surface. Some losses however, appear owing to the quality of the optical system.

As the concentration coefficient greater is, an accurate solar position tracking is necessary in order to preserve an efficient operation of the collector. This orientation must be permanent and needs an independent energy source.

The solar collectors with concentrating systems can be divided in two categories:
- collectors with motionless optical system
- collectors with mobile optical system

5.3.1 Collectors with motionless optical system

The flat collector associated with side mirrors

The simplest way to concentrate the solar radiation is to add one or two side mirrors to a flat collector (fig.12). The solar radiation arrives at the flat collector surface directly and reflected from the mirrors.

The concentration coefficient is variable, depending on the sun position on the sky. Sometimes, the entire power reflected from the mirrors reaches the flat collector surface. Other times, only a part of the reflected radiation arrives at the collector surface, depending on the incidence angle on the mirrors.

Flat collector associated with parabolic mirrors

Two segments of parabolic mirrors, $P_1$ and $P_2$ are arranged to lean on the side ends of a flat collector; the axes of the parabolic surfaces are disposed symmetrically against the flat collector axis, under an angle $\theta$. The both centers of the parabolic surfaces, points B and D, are located on the flat collector surface. The parabolic surfaces concentrate the incident rays towards the mirror center, if that rays are parallel with the mirror axe.
If the angle between the collector axe of symmetry and the incident rays are smaller than $\theta$, all the solar power entering the collector arrives at the absorbing surface. If that angle is greater than $\theta$ the solar energy is partially or totally reflected from a mirror towards another mirror and then back to the atmosphere. Depending on the size of angle $\theta$, the collector can operate 5-7 hours daily.

The geometrical concentration coefficient is

$$c = \frac{1}{\sin \theta}.$$

Actual values of this coefficient vary from 2 to 10.

**The spherical concentrator with mobile receiver**

A spherical mirror reflects the directional radiation towards a linear focus, parallel with the incident rays (fig.14). If the mirror is fixed, the receiver must perform a motion with two degrees of freedom in order to follow the sun path on the sky. The receiver movement needs smaller power consumption than the mirror orientation.

The concentration coefficient becomes highest if the mirror angle is $120^\circ$. For a perfect mirror surface, this maximum is 187.

Fig.14 The spherical concentrator

**5.3.2 Collectors with mobile concentrator**

**The parabolic mirror**

A parabolic mirror is obtained by rotating a parabola around its axis. This mirror has the highest concentration coefficient among all other types. Fig. 15 shows the optical behavior of this mirror.

The sun disc is seen from the sun-earth distance as the base of a circular cone having the top angle of 32'. The two side rays which arrive at the mirror top (point A) reflect maintaining the same angle between them, so that in the focal plane appears a circular light spot with diameter $d_f$.

The rays cone which arrives in point B, reflects, too and, crossing the focal plane, an ellipsoidal light spot appears, having the main axis $d > d_f$. Taking into account all the mirror points situated at angular distance $\theta$ from the mirror axe, the image in the focal plan becomes a circle with diameter $d$. If $\theta = 90$ degrees, all the focal plane will be illuminated and if $\theta > 90$ degrees, a concentrate light will appear on the surface exposed to the sun. These two last situations are not interesting for sun collectors because, obviously, the power receiver surface must be smaller than the mirror aperture in order to not shadow its surface.
Since as greater is the angle $\theta$, the focal spot diameter grows, the power reflected from the mirror surface spreads on a greater surface in the focal plane. Thus, the power density on the focal plane surface diminishes from center to edge. Theoretically, the power density changes as fig.16 shows; the curve 1 has a flat zone corresponding to the circle of diameter $d_f$. Because of deviations of the actual mirror from the geometrical shape, the power density changes as the curve 2 shows.

The diameter of the absorbing surface chooses according to the purpose that is intended for the mirror. That purpose can be:

- to obtain the higher possible temperature on the receiver;
- to collect the entire radiation reflected from the mirror surface.

In the first case, the receiver diameter must be equal with $d_f$. The maximum concentration coefficient is reached for an aperture angle of 90 degrees and is equal to

$$c_{\text{max}} = \frac{1}{\text{tg}^2 16'} \approx 46100.$$  \hspace{1cm} (24)
In the second situation the receiver surface chooses of diameter \( d \). The concentration coefficient is defined as

\[
c = \frac{S_D}{S_d} = \left( \frac{D}{d} \right)^2
\]  

(25)

We note, in fig. 17, \( AD = f \) and \( BF = R \). Considering the triangle BCF, results

\[
D = 2BC = 2R \sin \theta.
\]  

(26)

From the triangles GFN and FHM:

\[
d = MN = \frac{GH}{\cos \theta}
\]  

(27)

Since

\[GF = FH = R \tan 16',\]

results

\[
d = \frac{2R \tan 16'}{\cos \theta}
\]  

(28)

Finally

\[
c = \left( \frac{\sin \theta \cos \theta}{\tan 16'} \right)^2 = \left( \frac{\sin 2\theta}{2 \tan 16'} \right)^2,
\]  

(29)

that means the concentration coefficient depends on the aperture angle \( \theta \).

The function \( \sin 2\theta \) has the maximum value equal to 1, and that happens if \( 2\theta = 90 \) degrees, or \( \theta = 45 \) degrees.

For \( \theta = 45 \) degrees the concentration coefficient reaches the value

\[
c_{\text{max}} = \frac{1}{4 \tan^2 16'} \approx 11525.
\]  

(30)

In fact, the realized mirrors have deviations from the theoretical geometric shape so that the optimum aperture grows to 60 degrees. In this case the concentration coefficient can be only 8655. Even this value can not be obtained; practical values are 4-5 times smaller.

The cylindrical parabolic mirror is obtained by translating a parabola along a straight line. The sun image on the focal plane is now a rectangular one, which length is equal to the mirror length. The geometrical concentration coefficient is

\[
c = \frac{S_D}{S_d} = \frac{D}{d},
\]  

(31)

having the maximum value

\[
c_{\text{max}} = \frac{1}{2 \tan 16'} \approx 107.
\]  

Large parabolic mirrors are built from many small mirrors, so disposed to follow a parabolic surface. The small mirrors may be flat or curved. Such a mirror will have a smaller concentration coefficient, of course, than a parabolic mirror. Two examples are shown in fig. 18 and fig. 20.

5.4 Solar thermal power plant

Collecting the solar radiation with large mirror systems, a high steam temperature may be
obtained to generate electricity through the conventional thermodynamic cycle. Two main types of such plant are built.

a) **The plant with distributed collectors** (fig. 19).

Every collector has a mirror and a receiver where the heat carrier is warmed up. A pipe system collects the flows of heat carrier from all receivers and transports it to the steam generator. The size of such a plant is limited, up to some hundreds of kW, because the heat losses in the collecting pipe system.

b) **The plant with field of mirrors and central receiver** (fig. 20).

The receiver is placed on a tower top where the reflected radiation arrives from a very large number of mirrors arranged around the tower. The energy losses are much smaller in this case because the absorption of the solar radiation is very small in air and the length of the high temperature pipes are limited to the tower height. Some realized plants have installed power up to tens of MW.
Fig. 20 Parabolic system with a field of mirrors